

Blue Tides

On BlueWaters

“The first galaxies
and quasars”

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Rupert Croft (CMU)

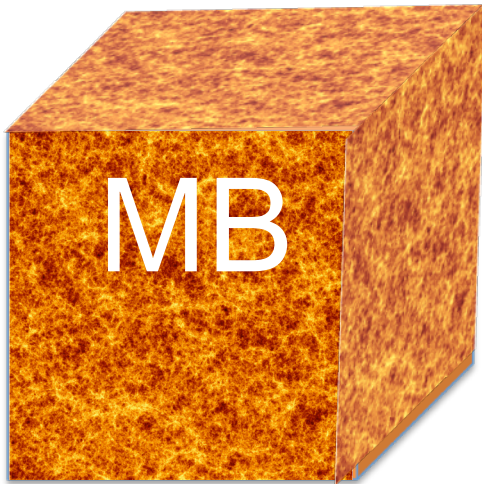
Nick Battaglia (Princeton)

Mark Straka (NCSA)

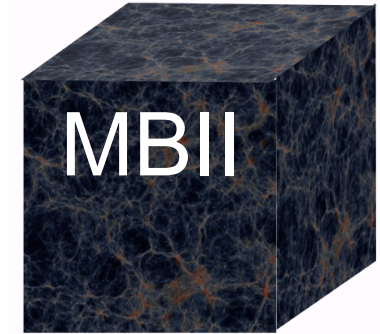
<http://bluetides-project.org>



Cosmological Hydro Simulations:



- Code used: **PetaGadget** (Petapps Cosmology)
- Physics: gravity, SPH, cooling, star formation, feedback, **black holes**.



- Particle number: $2 \times 3200^3 = 64$ billion
- Box size: $533 h^{-1}$ Mpc
- $Z_{\text{final}} = 4.75$

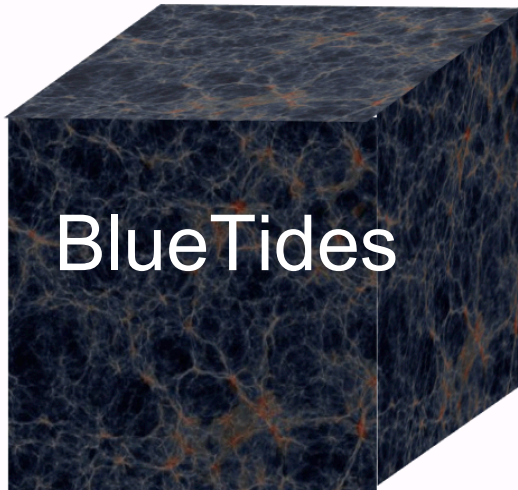
- $2 \times 1800^3 = 11.5$ billion
- $100 h^{-1}$ Mpc
- $Z_{\text{final}} = 0$ (biggest SPH vol)



- Runs using Kraken at NICS (>100k compute cores).

Team: N. Khandai, Y. Feng, C. DeGraf R. Croft, V. Springel, E. Tucker

Cosmological Hydro Simulations:



- Code used: **MPGadget** (Petapps Cosmology)
- Physics: p-SPH, H2 +cooling, star formation, feedback, **black holes**, **Patchy Reionization**.

WHOLE BW run

- Particle number: $2 \times 7040^3 = 0.7$ trillion
- Box size: $400 h^{-1}$ Mpc
- $Z_{\text{final}} = 7$ (8)

- Snapshots: 86 x (47 TB each)

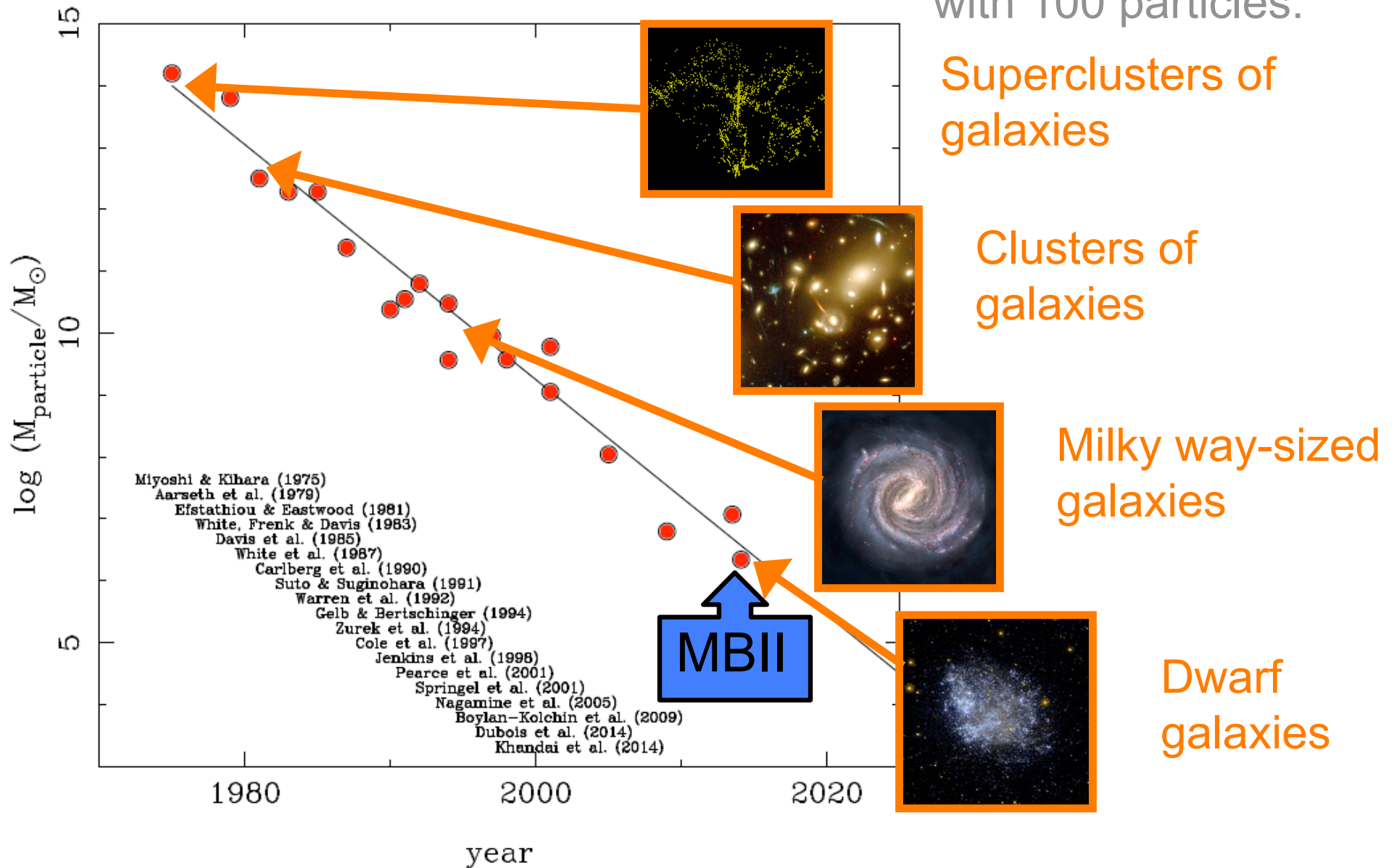


- Run using BlueWater at NCSA (648k compute cores).

Team: **Y. Feng**, DM, R. Croft, S. Bird, Battaglia

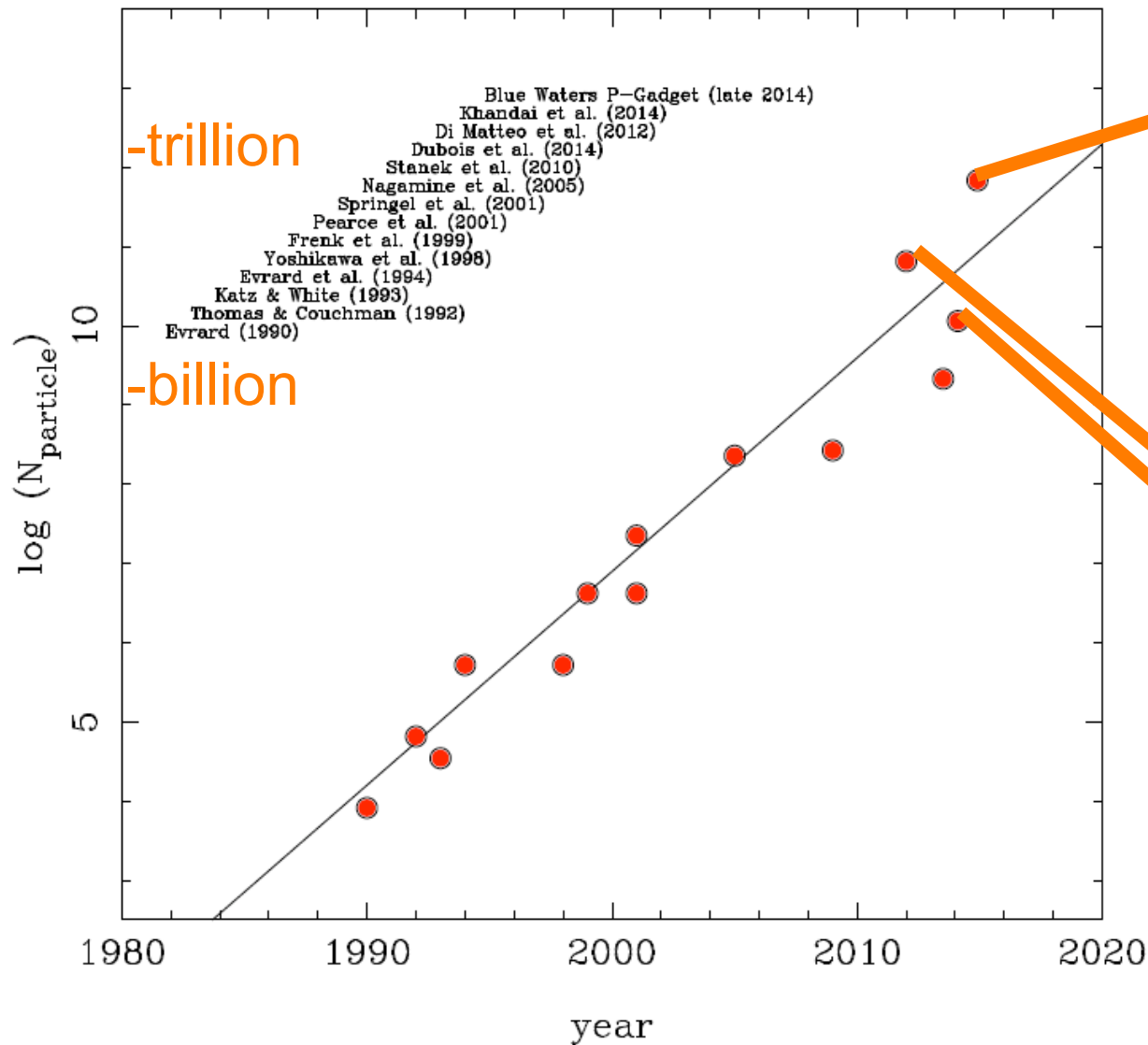
We resolve galaxies across the full mass function

What we can resolve with 100 particles:



Algorithms keep up with computational power

Hydro simulations:



On 0.72m cores
NCSA Cray XE6
Blue Waters

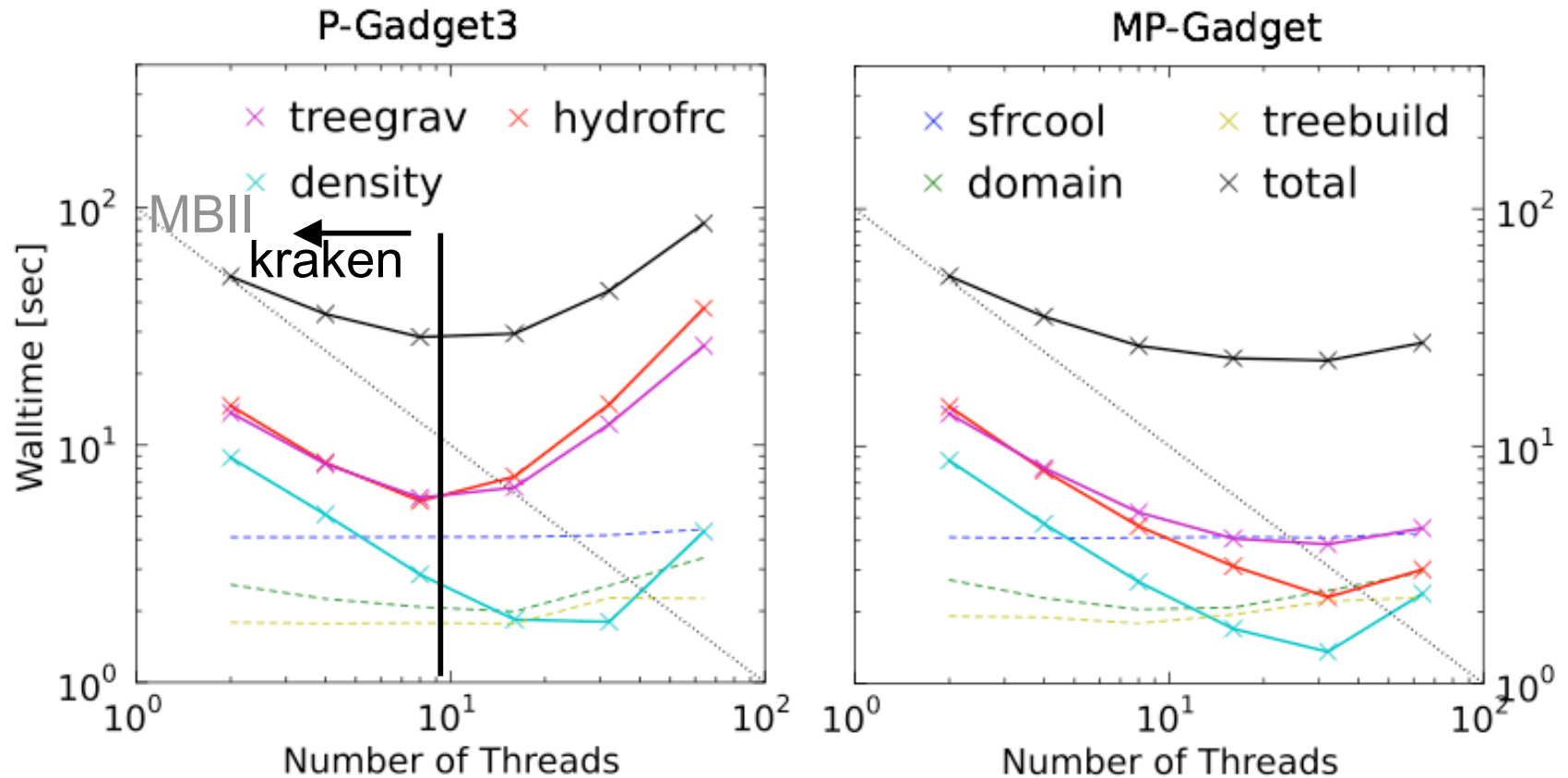


On 112k cores
NICS Cray XT5
Kraken

MP-Gadget:
Petascale cosmological code
(P-Gadget3)

Feng et al. 2015a,b, Code paper, in prep.

Short range force calculation: increased threading efficiency

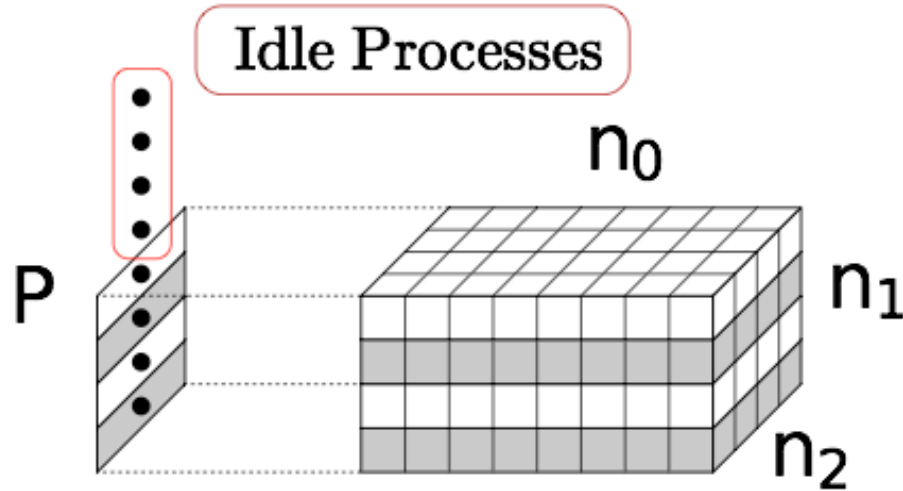


replaced global critical sections with spinlocks (per particle lock) and atomic increment operations

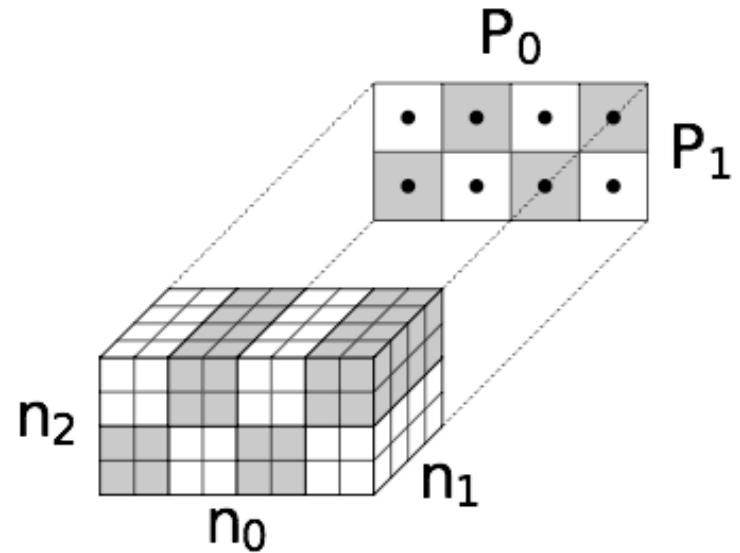
➔ 2 x speed-up

Long range force calculation (PM): New solver:

E.g. 8 processes:



FFTW/P-Gadget3



PFFT/MP-Gadget

Figure from M.Pippig 2013

Blue Tides:

N= 10000 slabs
on 81000 MPI ranks

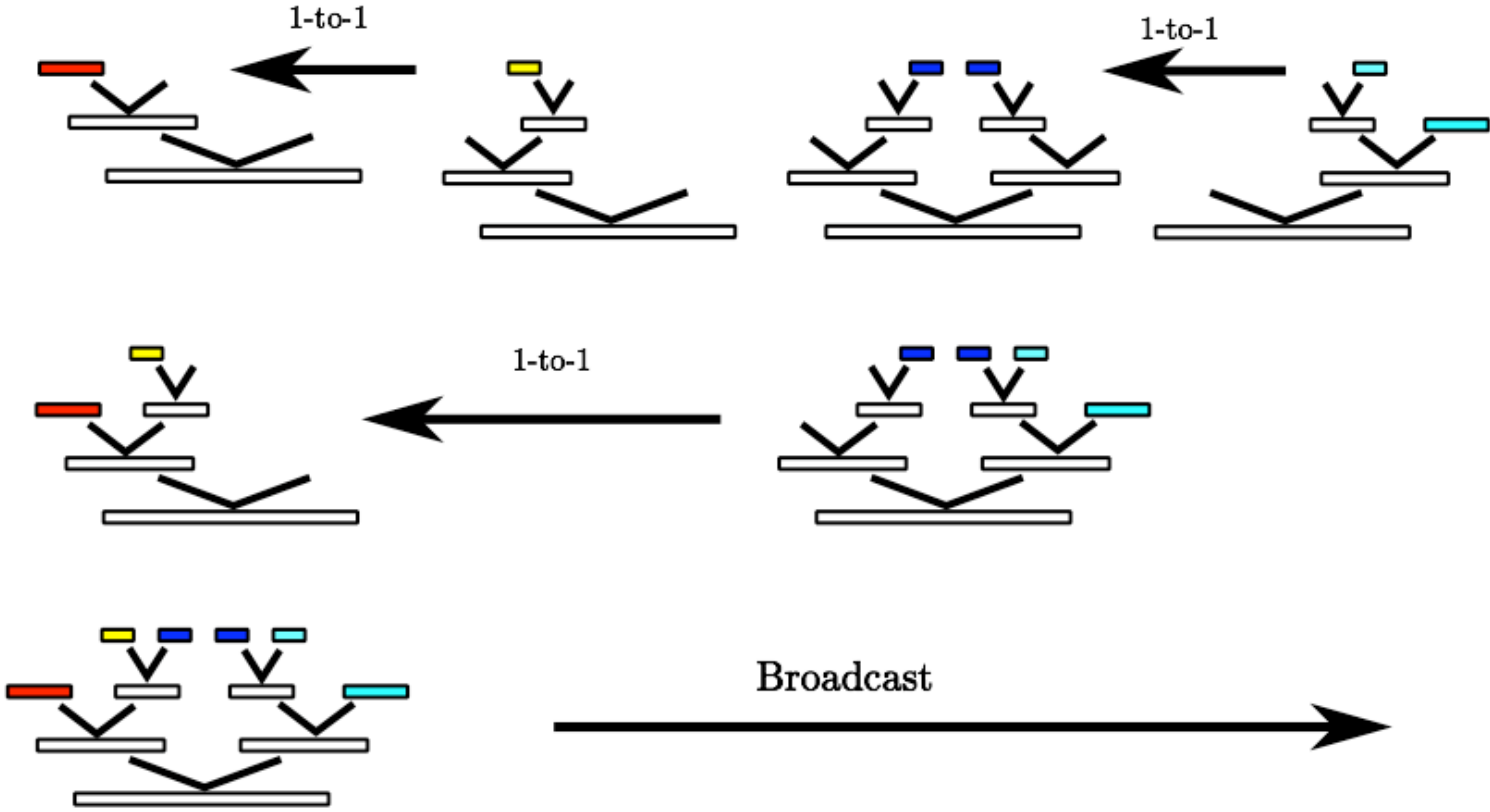


Pencil beam domain
decomposition

8 x speed-up

Open Source: Added new Array-execution interface and python binding to PFFT
(<http://github.com/mpip/pfft>)

Domain decomposition: **New global domain tree** built on root rank.



The communication is minimum, with one direction communication from children to root-ranks in each sub-communicator, followed by a global broadcast of the fully merged tree to all computing ranks.

10 x speed-up
(communication)

New parallel sorting module:

MP-Sort: histogram based Sorting: exchange 1 data item exactly once:

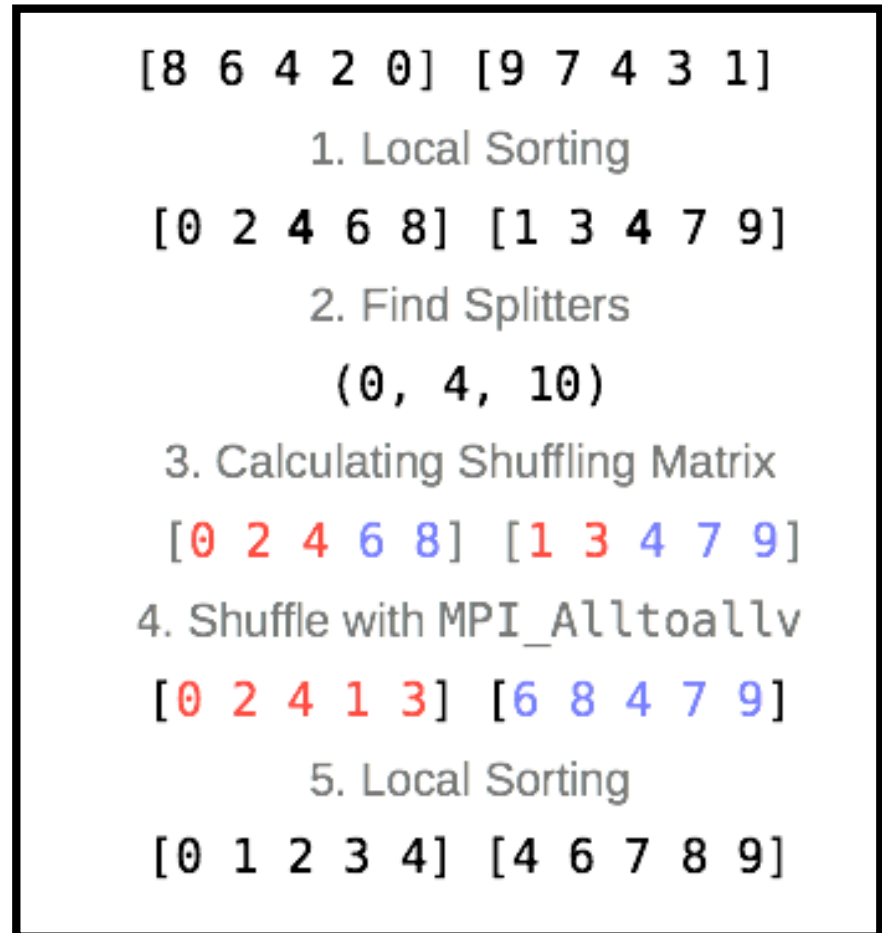
E.g sort 10 items on 2 MPI tasks

Large impact on IO,
FoF/galaxy catalogues

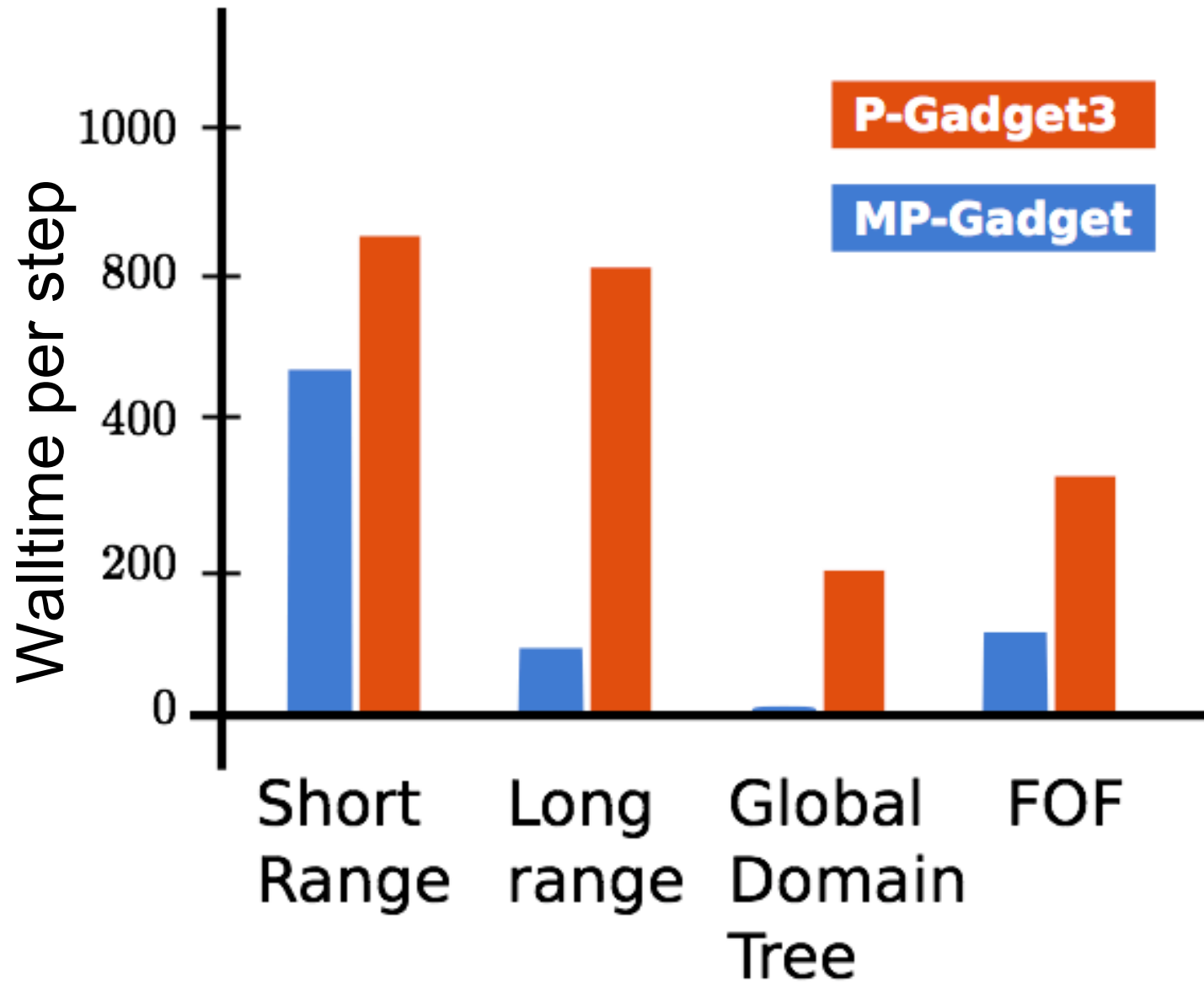
“Sorting At Scale on BlueWaters”
Y. Feng, M. Straka, R. Croft, TDM,
2015, CUG2015; Finalist of Best Paper.

Open Source:

<http://github.com/rainwoodman/MP-sort>



Code performance improvement for BlueTides



New Physical Modeling

☀ P-SPH formulation

☀ H₂ Molecular cooling/ star formation

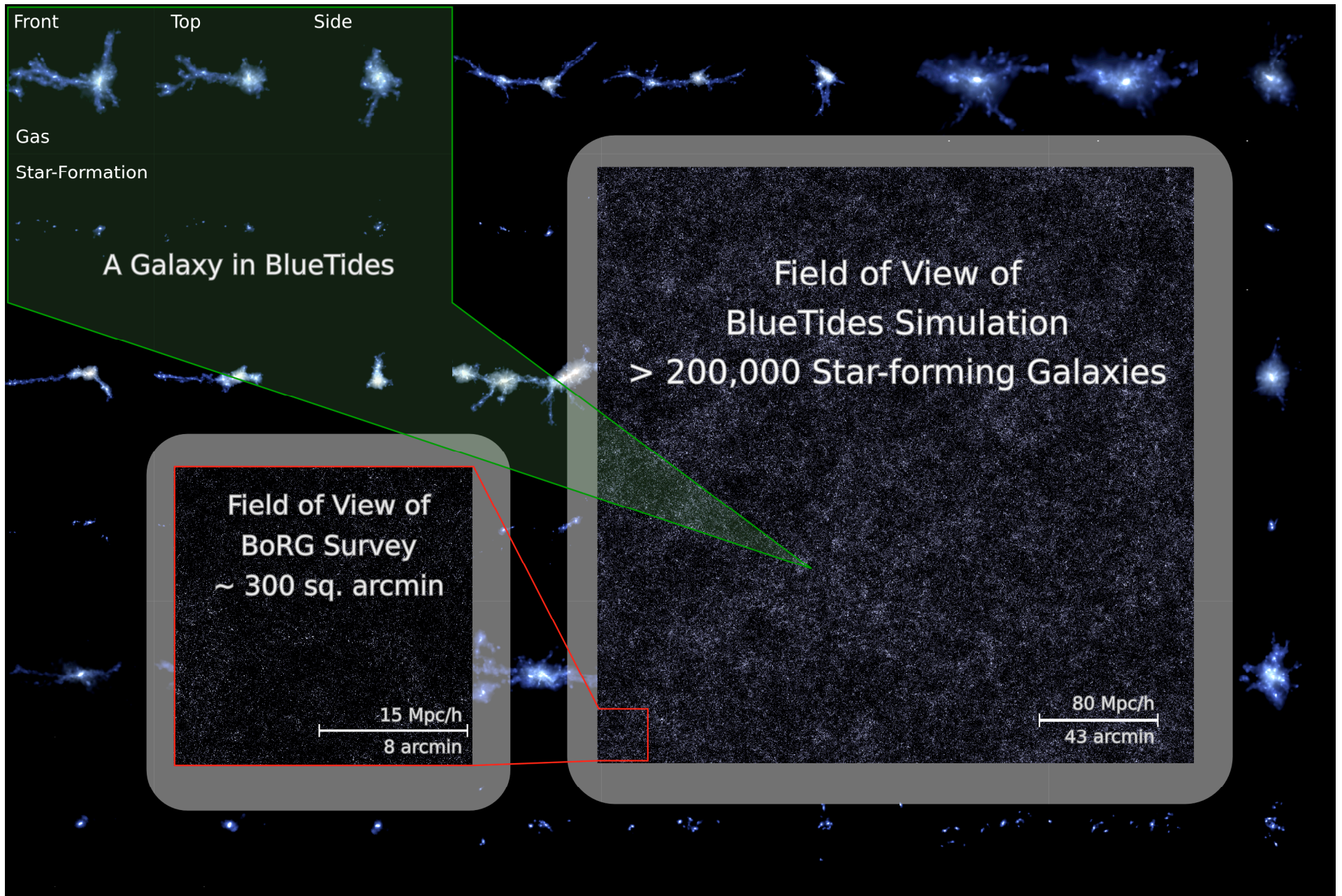
☀ Patchy Reionization (introducing spatially dependent
UV field, Battaglia et al. ,2013)

☀ Mass dependent Supernova Wind

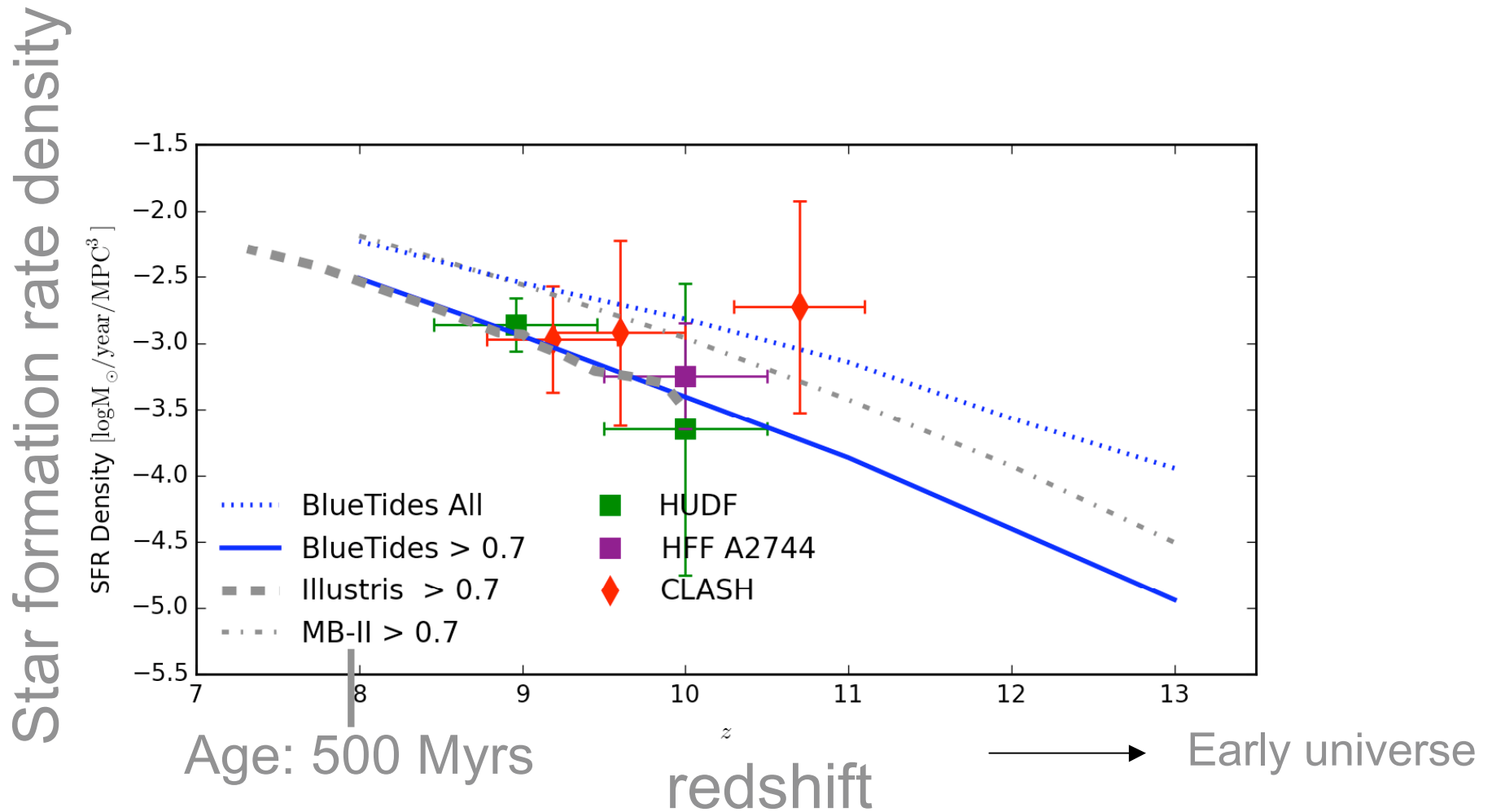
An example Problem :
What are the first galaxies
like?

Hubble Legacy Deep Fields: galaxies at $z=8-10$

Current Hubble Legacy Deep Fields probe tiny regions



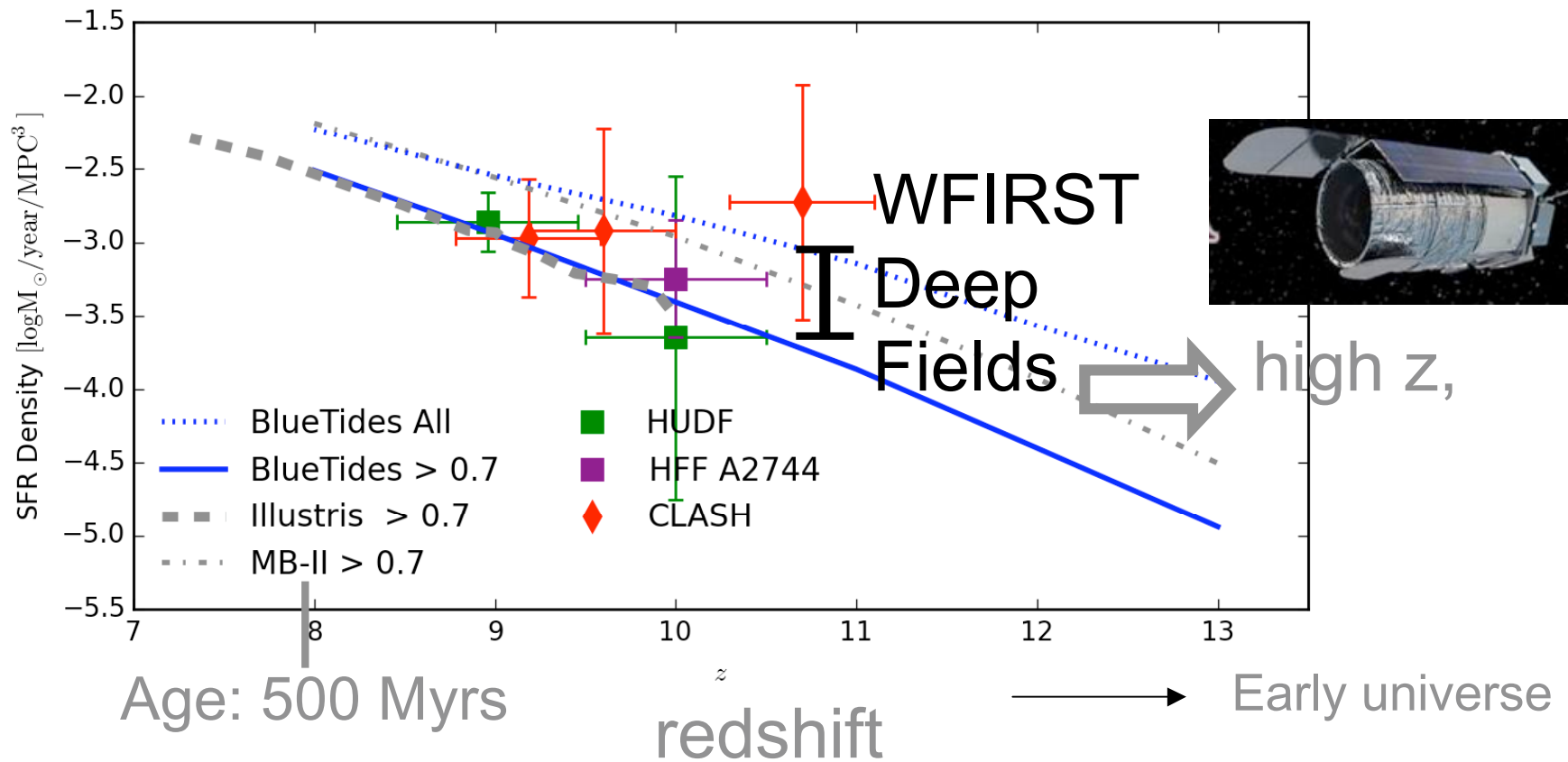
BlueTides Simulation: Global SFRD is consistent with current observations.



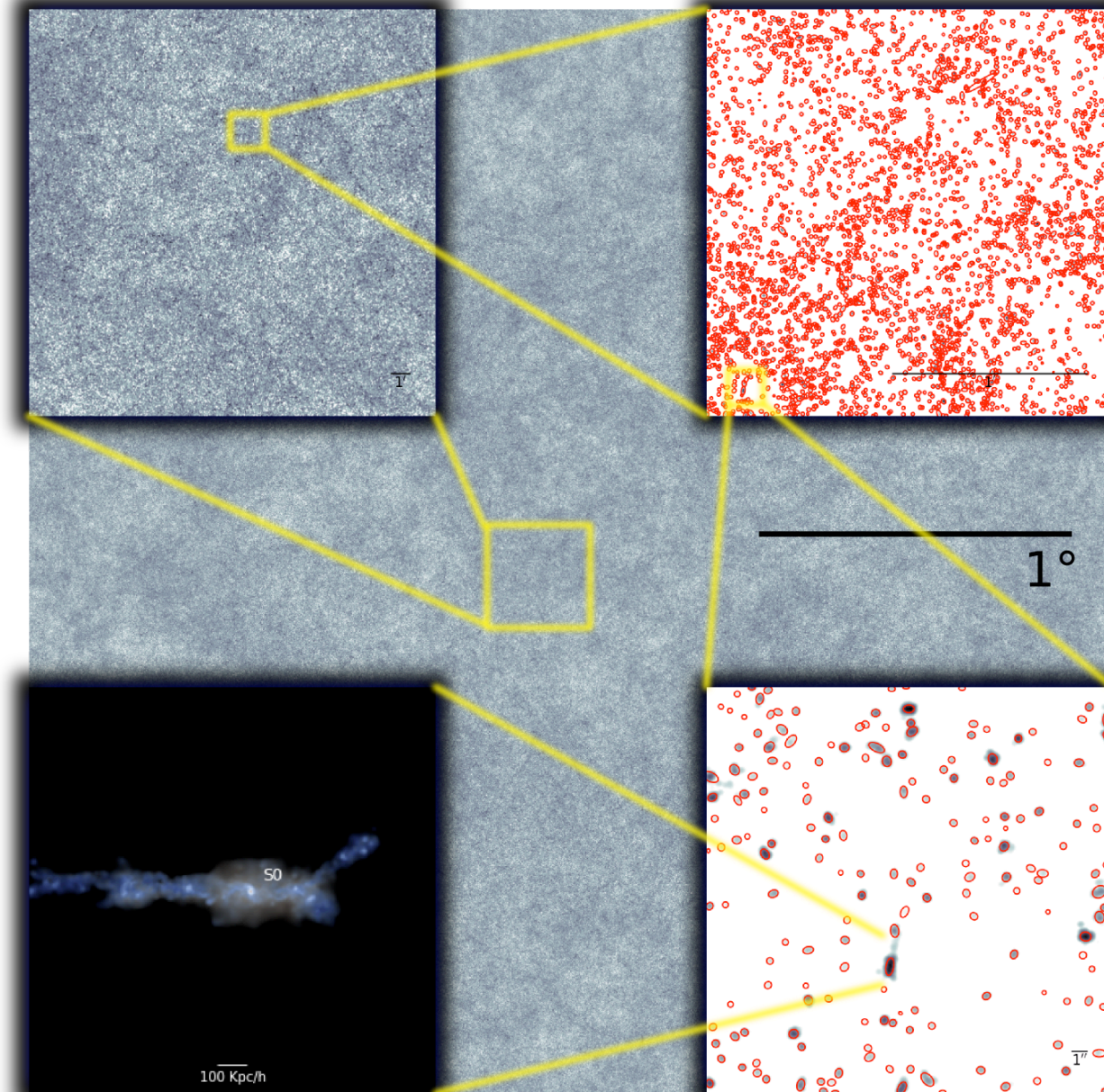
Feng et al., 2015a

BlueTides Simulation: Global SFRD is consistent with current observations.

Star formation rate density

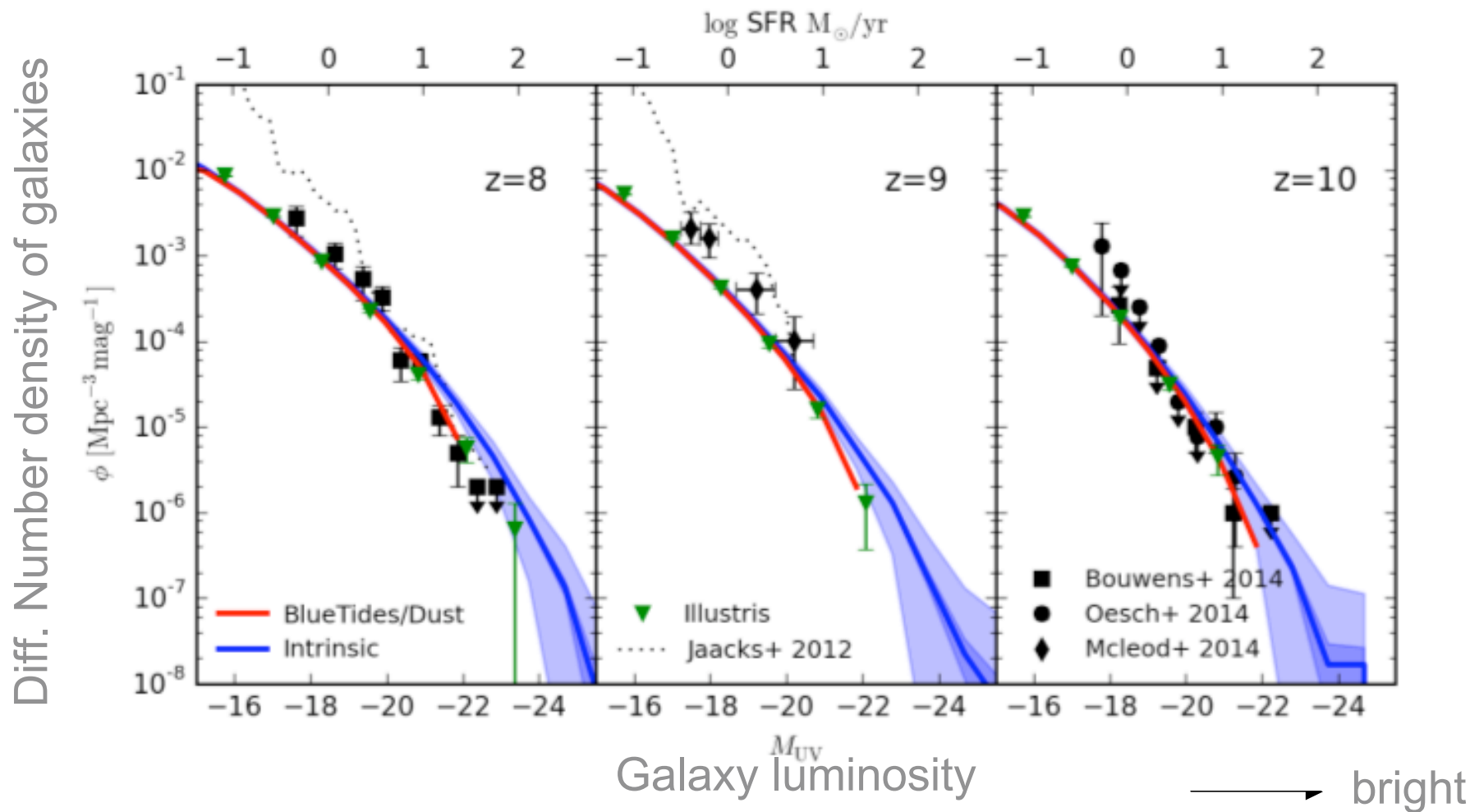


Simulations like Observations: Create Mock Fields.
Source extract detection to find galaxies



Galaxy Luminosity Function in BlueTides consistent with Hubble Legacy Fields

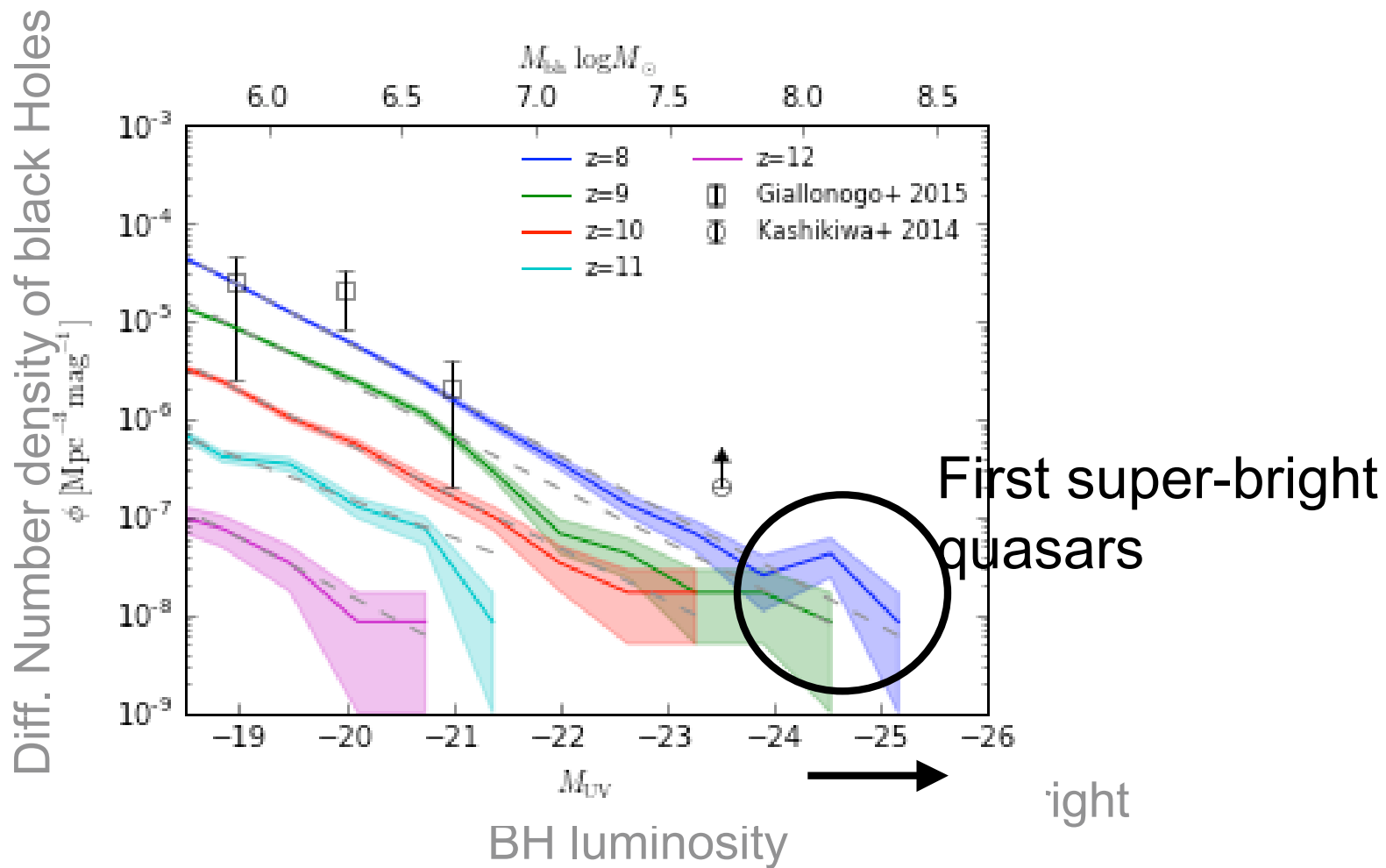
(star formation rate)



Feng et al., 2015a

Black Hole/AGN Luminosity Function in BlueTides

Consistent \longrightarrow Predictions for first quasars

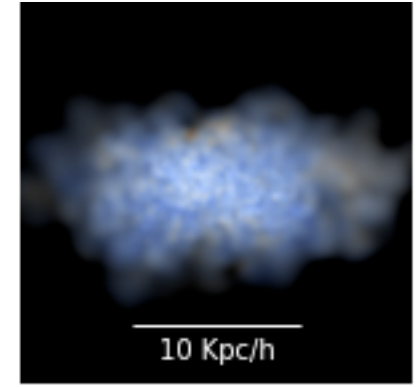
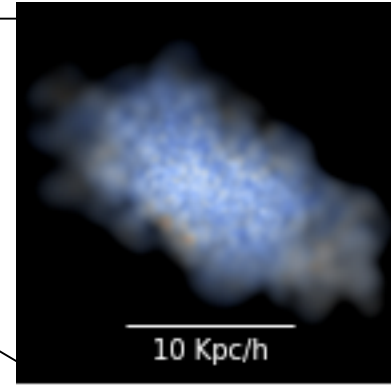
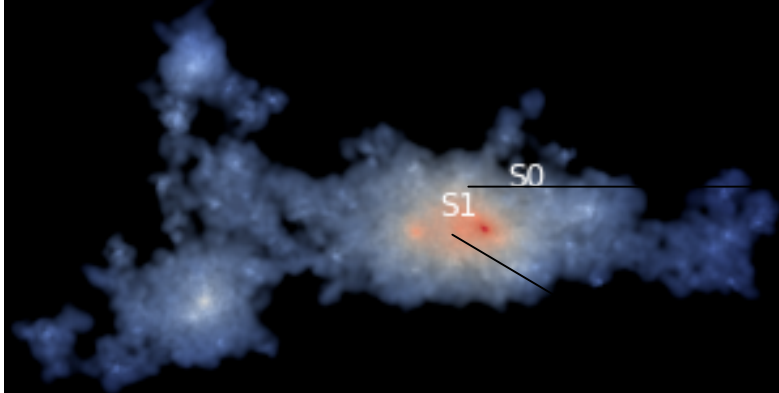


Feng et al., 2015a

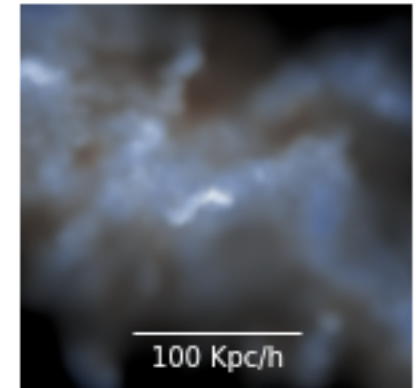
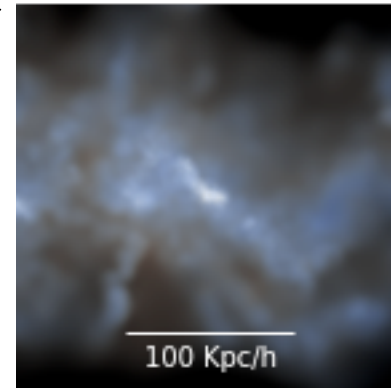
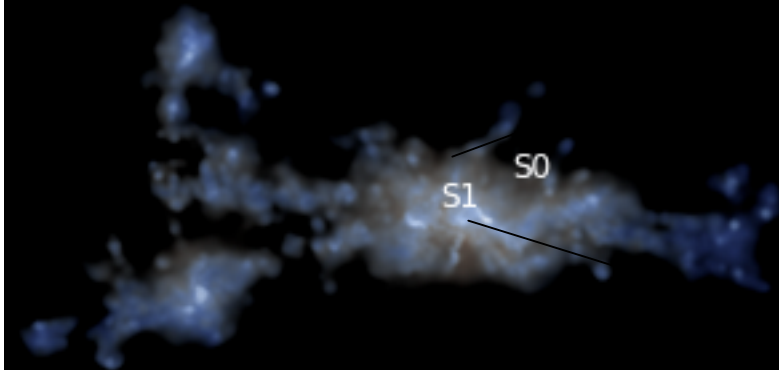
What are the first galaxies
like?

First galaxies are messy....

stars



gas



100 Kpc/h

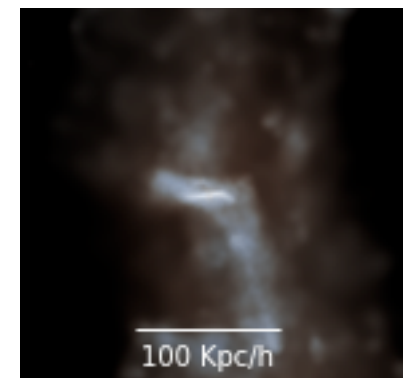
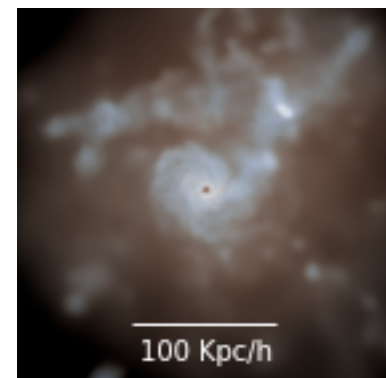
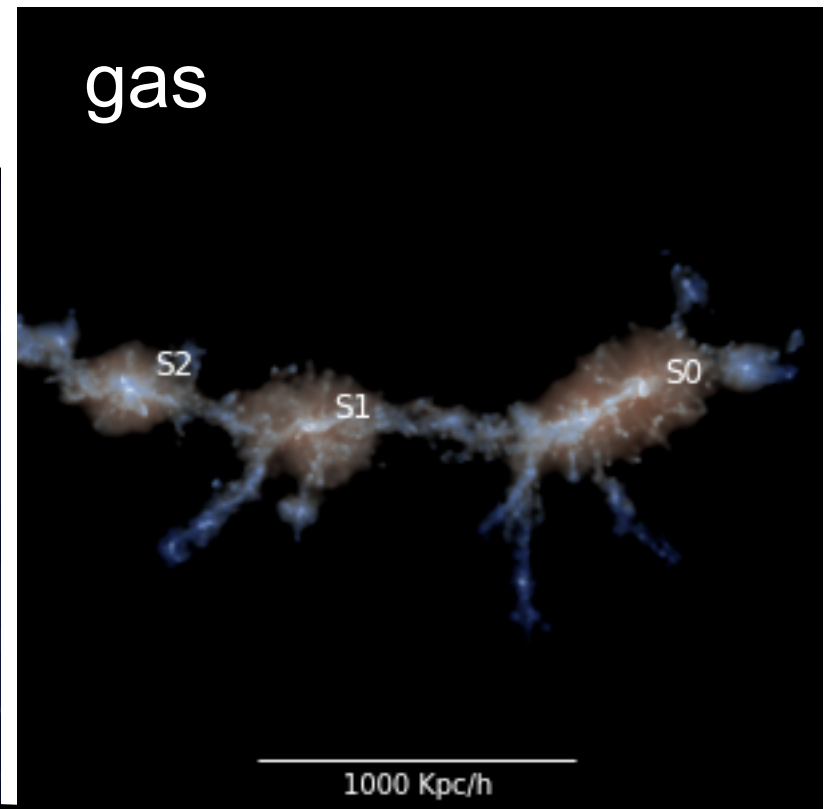
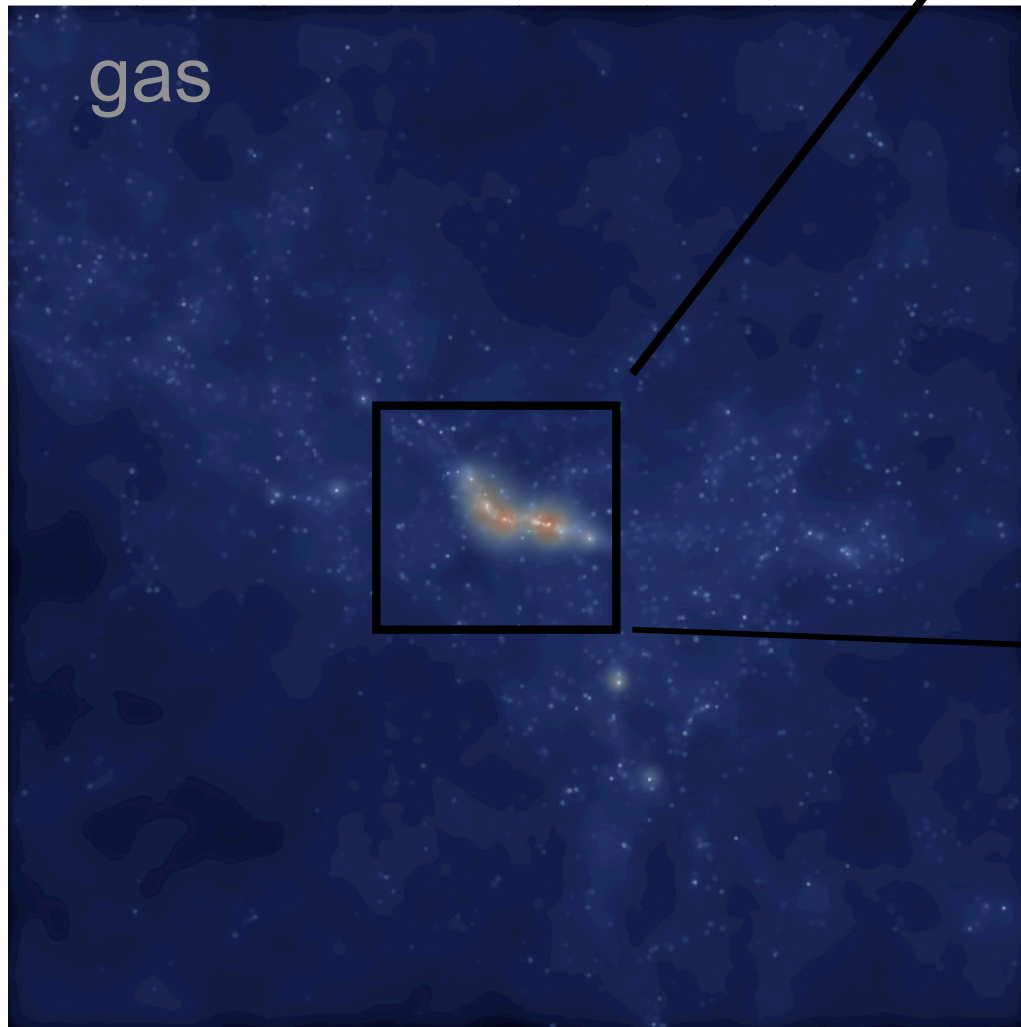
10 Kpc/h

10 Kpc/h

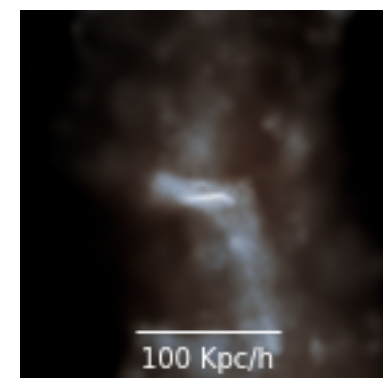
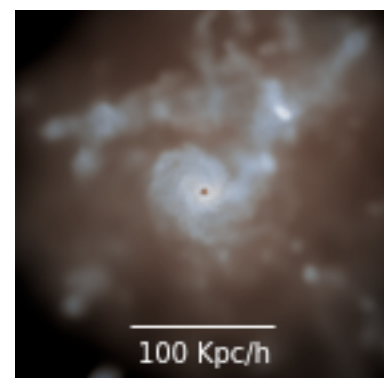
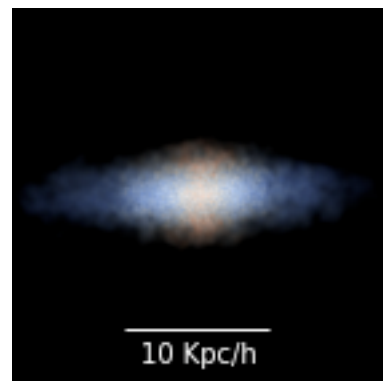
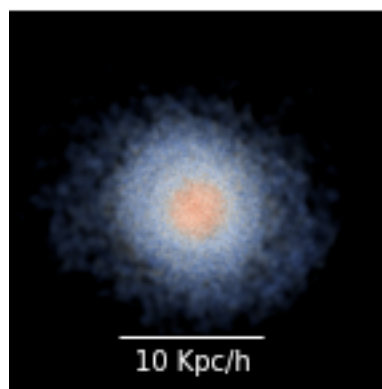
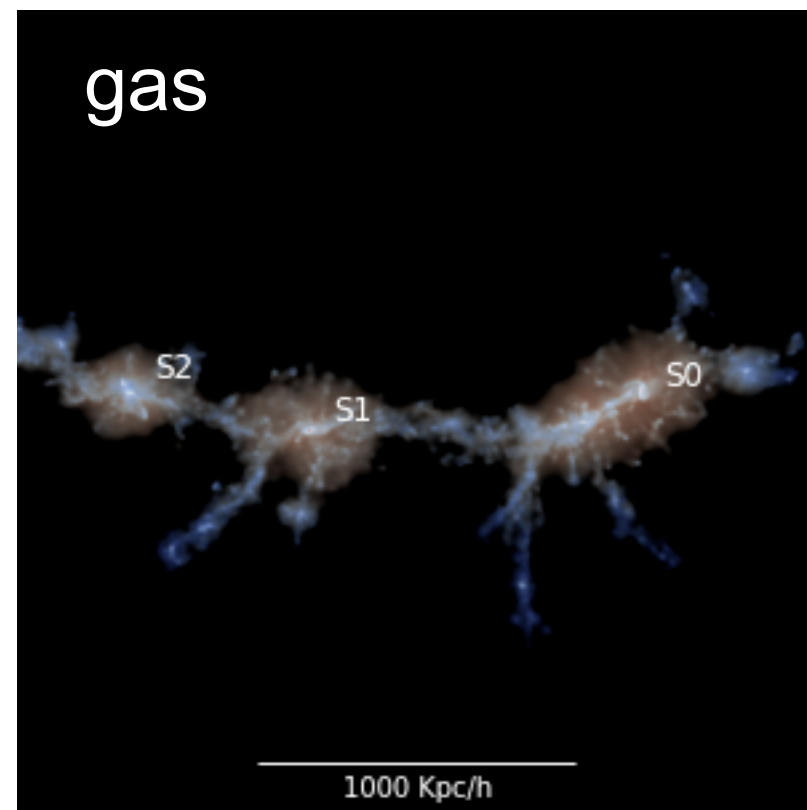
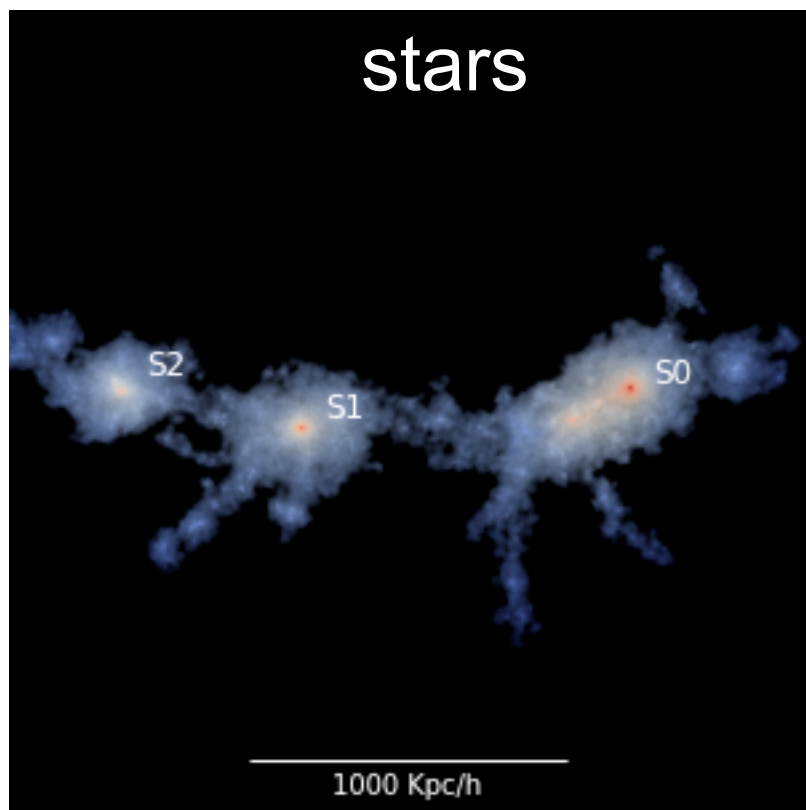
100 Kpc/h

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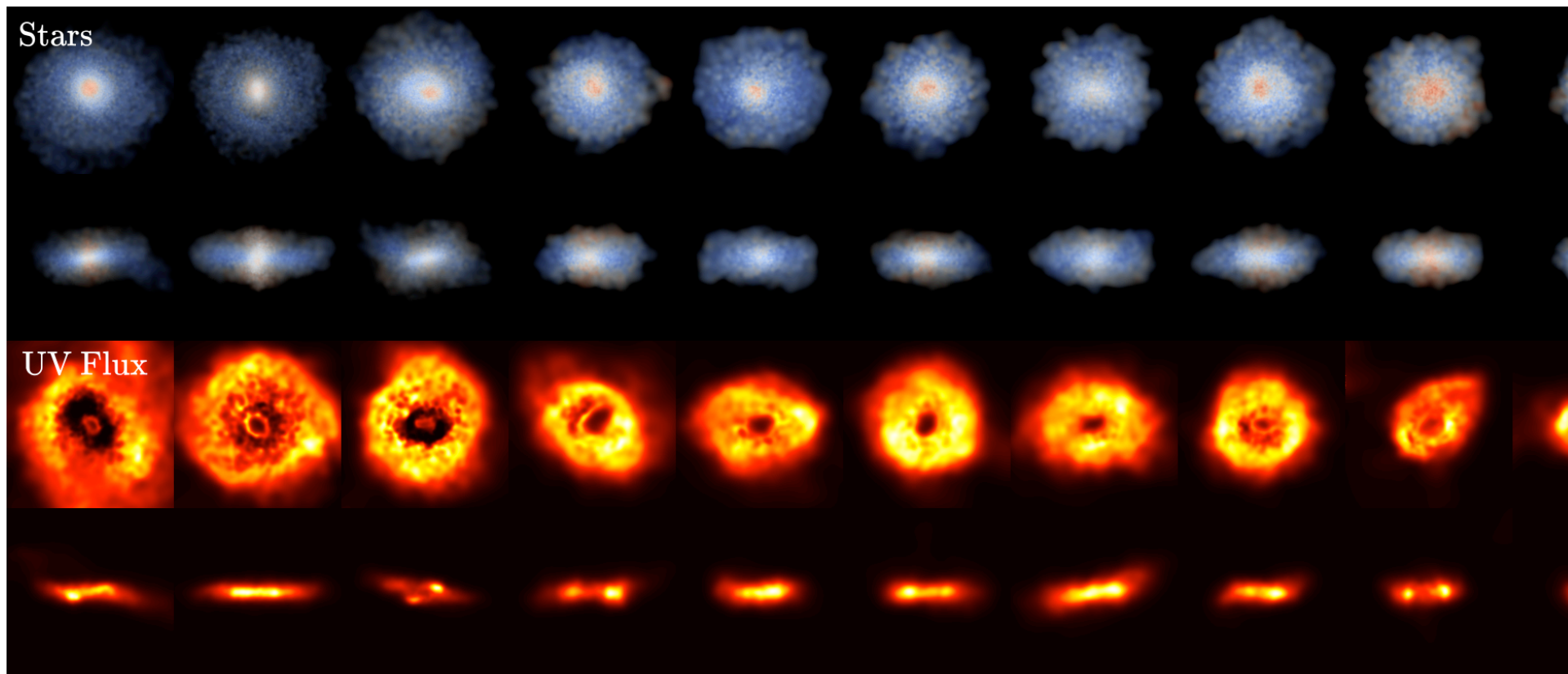
z=8 Most massive (Milky Way) galaxies



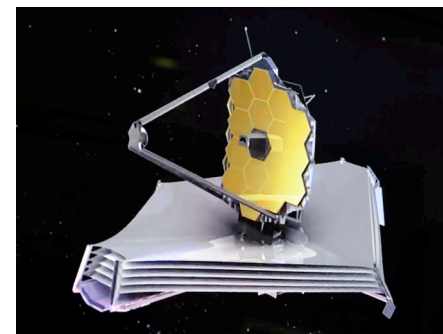
$z=8$ Most massive (Milky Way) galaxies are disks!



$z=8$ Milky Way (/Massive) Halos look like **disks!**



JWST



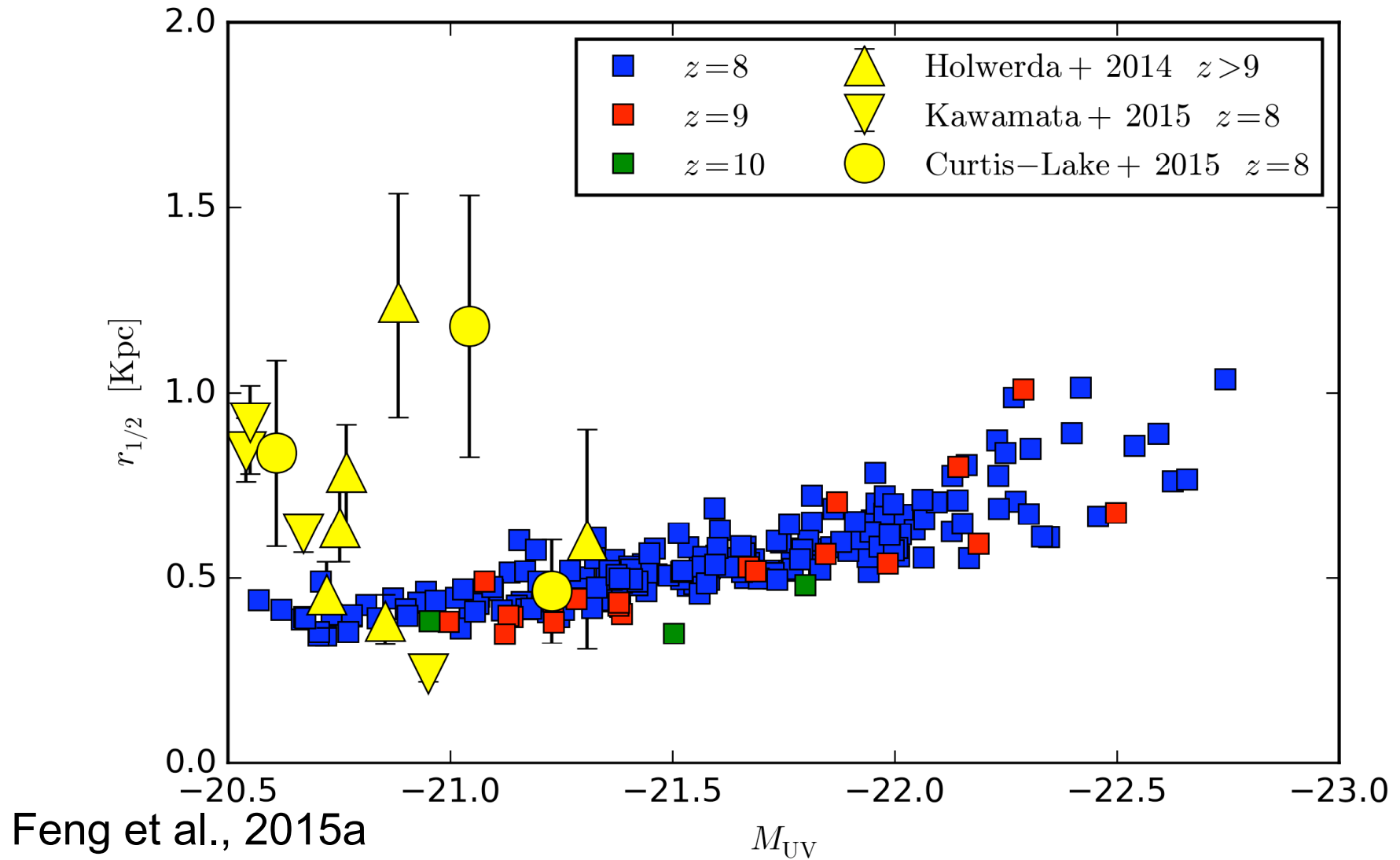
Feng et al., 2015b

Star formation in BlueTides (subgrid):

- Multiphase ISM (Springel 2003)
- SFR depends on Metallicity/Molecular Hydrogen
(Krumholz & Gnedin 2011)
- Supernova Feedback/wind depends on Halo Mass
(e.g. Okamoto 2010)

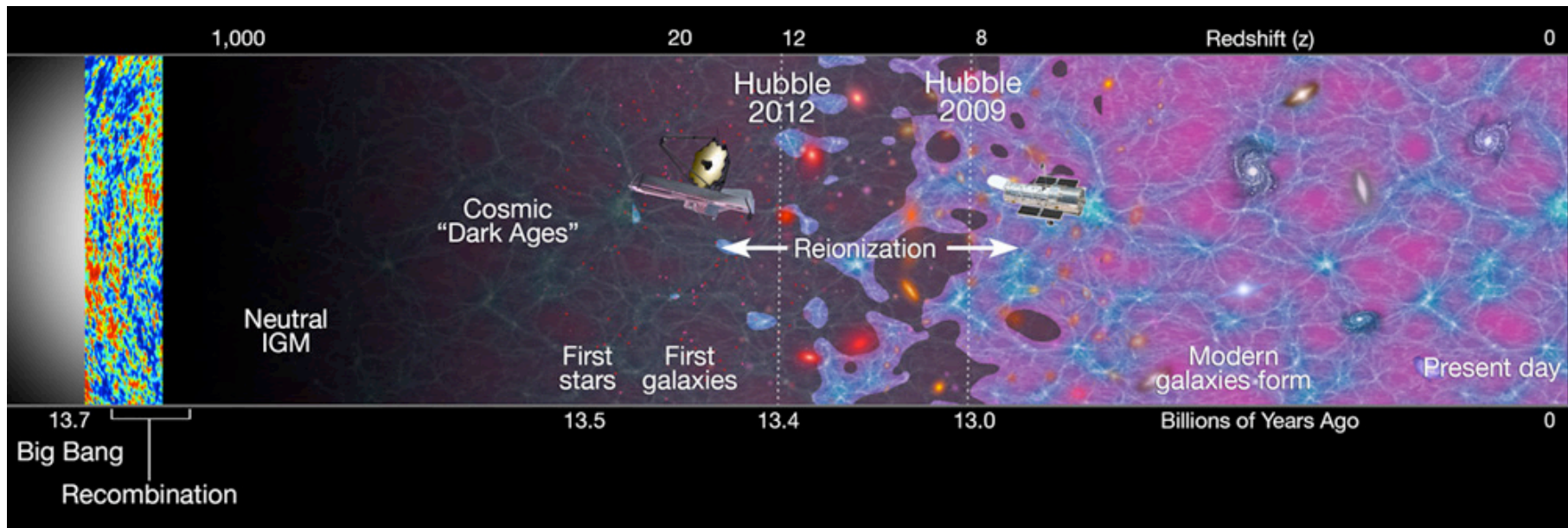
BH subgrid model as before

The sizes of galaxies in BlueTides are consistent with HST observations --> larger disks in bright galaxies



What sources reionize the Universe?

Galaxies and AGNs in BlueTides



BlueTides and Re-ionization history of the Universe

Galaxies can reionize the universe for high escape photon fractions. But AGNs can contribute (very?) significantly

